## Can EPR correlations be driven by and effective wormhole?\*

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## Abstract

We consider the two-particle wave function of an EPR system given by a two dimensional relativistic scalar field model. The Bohm-de Broglie interpretation is applied and the quantum potential is viewed as modifying the Minkowski geometry. In such a way singularities appear in the metric, opening the possibility, following Holland, of interpreting the EPR correlations as originated by a wormhole effective geometry, through which physical signals can propagate.

PACS numbers: 03.65.Ta, 03.65.Ud, 03.70.+k

<sup>\*</sup> Talk given at the Eleventh Marcel Grossmann Meeting, Berlin, Germany, 23-29 July 2006.

A causal approach to the Einstein-Podolsky-Rosen (EPR) problem, i.e. a two-particle correlated system, is developed. We attack the problem from the point of view of quantum field theory considering the two-particle function for a scalar field and interpreting it according to the Bohm - de Broglie view. In this approach it is possible to interpret the quantum effects as modifying the geometry in such a way that the scalar particles see an effective geometry. For a two-dimensional static EPR model we are able to show that quantum effects introduces singularities in the metric, a key ingredient of a bridge construction or wormhole. Following a suggestion by Holland [1] this open the possibility of interpret the EPR correlations as driven by an effective wormhole<sup>1</sup>.

The two-particle wave function of a scalar field,  $\psi_2(\mathbf{x_1}, \mathbf{x_2}, \mathbf{t})$  satisfies (see for example [3] [4]):

$$\sum_{j=1}^{2} [(\partial^{\mu} \partial_{\mu})_{j} + m^{2}] \psi_{2}(\vec{\mathbf{x}}^{(2)}, t) = 0$$
 (1)

where  $\vec{\mathbf{x}}^{(n)} \equiv \{\mathbf{x_1}, ... \mathbf{x_n}\}$ . Explicitly we have

$$[(\partial^{\mu}\partial_{\mu})_{1} + m^{2}]\psi_{2}(\mathbf{x}_{1}, \mathbf{x}_{2}, \mathbf{t}) + [(\partial^{\mu}\partial_{\mu})_{2} + \mathbf{m}^{2}]\psi_{2}(\mathbf{x}_{1}, \mathbf{x}_{2}, \mathbf{t}) = \mathbf{0}.$$

$$(2)$$

Substituting  $\psi_2 = R \exp(iS/\hbar)$  in Eq. (2) we obtain two equations, one of them for the real part and the other for the imaginary part. The first equation reads

$$\eta^{\mu_1\nu_1}\partial_{\mu_1}S\partial_{\nu_1}S + \eta^{\mu_2\nu_2}\partial_{\mu_2}S\partial_{\nu_2}S = 2\mathcal{M}^2 \tag{3}$$

where

$$\mathcal{M}^2 \equiv m^2 \hbar^2 (1 - \frac{Q}{2m^2 \hbar^2}) \tag{4}$$

with 
$$Q \equiv Q_1 + Q_2$$
 being  $Q_1 = -\hbar^2 \frac{(\partial^{\mu} \partial_{\mu})_1 R}{R}$  and  $Q_2 = -\hbar^2 \frac{(\partial^{\mu} \partial_{\mu})_2 R}{R}$ . (4')

The equation that comes from the imaginary part is

$$\eta^{\mu_1\nu_1}\partial_{\mu_1}(R^2\partial_{\nu_1}S) + \eta^{\mu_2\nu_2}\partial_{\mu_2}(R^2\partial_{\nu_2}S) = 0$$
 (5)

which is a continuity equation.

<sup>&</sup>lt;sup>1</sup> An extended version of this talk can be found in [2] where a non-tachyonic EPR model is studied

Following De Broglie [5] we rewrite the Hamilton-Jacobi equation (3) as

$$\frac{\eta^{\mu_1\nu_1}}{(1 - \frac{Q}{2m\hbar^2})} \partial_{\mu_1} S \partial_{\nu_1} S + \frac{\eta^{\mu_2\nu_2}}{(1 - \frac{Q}{2m\hbar^2})} \partial_{\mu_2} S \partial_{\nu_2} S = 2m^2\hbar^2.$$
 (6)

Here  $\eta^{\mu\nu}$  is the Minkowski metric and we can interpret the quantum effects as realizing a conformal transformation of the metric in such a way that the effective metric is  $g_{\mu\nu} = (1 - \frac{Q}{2m^2\hbar^2})\eta_{\mu\nu}$ . Now, following an approach by Alves (see [7]), we will see that for the static case it is possible to obtain a solution as an effective metric which comes from Eqs. (3) and (5). For the static case these equations are:

$$\eta^{11}\partial_{x_1}S\partial_{x_1}S + \eta^{11}\partial_{x_2}S\partial_{x_2}S = 2m^2\hbar^2(1 - \frac{Q}{2m\hbar^2})$$
 (7)

$$\partial_{x_1}(R^2\partial_{x_1}S) + \partial_{x_2}(R^2\partial_{x_2}S) = 0 \tag{8}$$

We consider that our two-particle system satisfies the EPR condition  $p_1 = -p_2$  which in the BdB interpretation, using the Bohm guidance equation  $p = \partial_x S$ , can be written as  $\partial_{x_1} S = -\partial_{x_2} S$ . Using this condition in Eq. (8) we have  $\partial_{x_1} (R^2 \partial_{x_1} S) = \partial_{x_2} (R^2 \partial_{x_1} S)$  and this equation has the solution  $R^2 \frac{\partial S}{\partial x_1} = G(x_1 + x_2)$  where G is an arbitrary (well behaved) function of  $x_1 + x_2$ . Substituting in Eq.(7) we have

$$2m^2\hbar^2(1 - \frac{Q}{2m\hbar^2}) = 2(\frac{G}{R^2})^2 \tag{9}$$

and using the expression (4') for the quantum potential, the last equation reads

$$8G^{2} + (\partial_{x_{1}}(R^{2}))^{2} - 2R^{2}\partial_{x_{1}}^{2}R^{2} + (\partial_{x_{2}}(R^{2}))^{2} - 2R^{2}\partial_{x_{2}}^{2}R^{2} - 8m^{2}R^{4} = 0.$$
 (10)

A solution of this nonlinear equation is  $R^4 = \frac{1}{2m^2}(C_1\sin(m(x_1+x_2)+C_2))$  provided an adequated function  $G(x_1+x_2)$  which can be obtained from (10) by substituting the solution.

In order to interpret the effect of the quantum potential we can re-write Eq. (7) using (9) obtaining  $m^2 \frac{\eta^{11}}{(\frac{G}{R^2})^2} \partial_{x_1} S \partial_{x_1} S + m^2 \frac{\eta^{11}}{(\frac{G}{R^2})^2} \partial_{x_2} S \partial_{x_2} S = 2m^2$  that we write as

$$g^{11}\partial_{x_1}S\partial_{x_1}S + g^{11}\partial_{x_2}S\partial_{x_2}S = 2m^2$$

$$\tag{11}$$

and then we see that the quantum potential was "'absorbed"' in the new metric  $g_{11}$  which is:

$$g_{11} = \frac{1}{g^{11}} = \frac{\eta_{11}}{m^2} \left(\frac{G}{R^2}\right)^2 = \frac{-\frac{C_1^2}{16m^2} + \frac{3C_1^2}{16m^2} \sin^2(m(x_1 + x_2) + C_2)}{\frac{C_1}{2m^2} \sin(m(x_1 + x_2) + C_2)}.$$
 (12)

We can see that this metric is singular at the zeroes of the denominator in (12) and this is characteristic of a two dimensional black hole solution (see [6] [7]). Then our two-particle system "see" an effective metric with singularities, a fundamental component of a whormhole[8]. This open the possibility, following Holland [1], of interpret the EPR correlations of the entangled particles as driven by an effective wormhole. Obviously a more realistic (i.e. four dimensional) and more sophisticated model (i.e. including the spin of the particles) must be studied. <sup>2</sup>

## Acknowledgements

I would like to thank Prof. Nelson Pinto-Neto, from ICRA/CBPF, Prof. Sebastião Alves Dias, from LAFEX/CBPF, Prof. Marcelo Alves, from IF/UFRJ, and the 'Pequeno Seminario' of ICRA/CBPF for useful discussions. I would also like to thank Ministério da Ciência e Tecnologia/ CNEN and CBPF of Brazil for financial support.

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$$ds^2 = \Omega^2 dx^2 \tag{13}$$

with

$$\Omega^2 = 1 + \frac{b^2}{(x - x_0)^2}. (14)$$

This looks like a metric with a singularity at  $x_0$ . However, the divergence of the conformal factor can be though as the space opening out to another asymptotically flat region connected with the first one by mean a wormhole of size 2b.

<sup>&</sup>lt;sup>2</sup> It is interesting to note that a wormhole coming from a (Euclidean ) conformally flat metric with singularities was shown by Hawking [9]. Consider the metric:

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